

# Search for large extra dimensions in the mono-photon final state at $\sqrt{s} = 1.96$ TeV

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We report on a search for large extra dimensions in a data sample of approximately  $1 \text{ fb}^{-1}$  of  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96 \text{ TeV}$ . We investigate Kaluza-Klein graviton production with a photon and missing transverse energy in the final state. At the 95% C.L. we set limits on the fundamental mass scale  $M_D$  from 884 GeV to 778 GeV for 2 to 8 extra dimensions.

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Arkani-Hamed, Dimopoulos, and Dvali (ADD) [1] made the first attempt to solve the hierarchy problem of the standard model (SM) by postulating the existence of  $n$  new large extra spatial dimensions (LED). In this approach, the SM particles are confined to a 3-dimensional brane while gravity is diluted in the larger volume. The size of the compactified extra space ( $R$ ), the effective Planck scale in the 4-dimensional space-time ( $M_{Pl}$ ), and the fundamental Planck scale in the  $(4+n)$ -dimensional space-time ( $M_D$ ), are related by the expression  $M_{Pl}^2 = 8\pi M_D^{n+2} R^n$ . Due to the compactification of the extra space, the gravitational field appears as a series of quantized energy states which are referred as Kaluza-Klein modes with mass splittings  $\Delta m \approx 1/R$ . For a moderate number of extra dimensions ( $n \leq 8$ ), the mass splitting is small enough that the different modes can be integrated. A Kaluza-Klein graviton ( $G_{KK}$ ) behaves like a massive, non-interacting, stable particle whose direct production gives an imbalance in the final state momentum as its collider signature.

In this Letter we report the results of a search for LED in  $p\bar{p}$  collisions at  $\sqrt{s} = 1.96$  TeV, using the D0 detector at the Fermilab Tevatron collider in the exclusive single photon plus missing transverse energy ( $\gamma + \cancel{E}_T$ ) final state. This signature arises from the process  $q\bar{q} \rightarrow \gamma G_{KK}$  which is studied in detail in [2]. The CDF experiment carried out a similar search with 87 pb<sup>-1</sup> of data, setting 95% C.L. lower limits on the fundamental Planck scale  $M_D$  of 549, 581 and 601 GeV for 4, 6, and 8 extra dimensions, respectively [3]. Several other searches for LED have been performed by collaborations at the Tevatron [4, 5] and the CERN LEP collider [6].

The backgrounds to the  $\gamma + \cancel{E}_T$  signal are dominated by electroweak boson production and non-beam collision background where muons from the beam halo or cosmic rays undergo bremsstrahlung producing an energetic photon. The former source is dominated by the  $Z + \gamma \rightarrow \nu\bar{\nu} + \gamma$  process, followed by  $W \rightarrow e\nu$  where the electron is misidentified as a photon,  $W + \gamma$  where the lepton from the  $W$  boson decay is not detected, and  $W/Z + \text{jet}$  production where the jet is misidentified as a photon.

The D0 detector [7] comprises a central-tracking system with a silicon microstrip tracker (SMT) and a central fiber tracker (CFT), both housed within a 2 T superconducting solenoidal magnet, with designs optimized for tracking and vertexing at  $|\eta| < 3$  and  $|\eta| < 2.5$ , respectively, where  $\eta$  is the pseudorapidity [8] measured with respect to the geometrical center of the detector. The central preshower system (CPS) is located in front of a liquid-argon/uranium calorimeter and consists of three layers of scintillating strips, providing precise measurement of EM shower positions. The calorimeter has a central section (CC) covering  $|\eta| \leq 1.1$ , and two end calorimeters (EC) that extend coverage to  $|\eta| \approx 4.2$ , each of them located in a separate cryostat [9]. Each

calorimeter contains an electromagnetic (EM) section closest to the interaction region followed by fine and coarse hadronic sections. The electromagnetic part has four longitudinal layers and transverse segmentation of  $0.1 \times 0.1$  in  $\eta - \phi$  space (where  $\phi$  is the azimuthal angle), with the exception of the third layer, where it is  $0.05 \times 0.05$ . Additionally, scintillators between the CC and EC cryostats provide sampling of developing showers for  $1.1 < |\eta| < 1.4$ . The outer muon system, covering  $|\eta| < 2$ , consists of a layer of tracking detectors and scintillation trigger counters in front of 1.8 T iron toroids, followed by two similar layers after the toroids. The data in this analysis were recorded using single electromagnetic object triggers that are almost 100% efficient to select signal events. It corresponds to an integrated luminosity of  $1.05 \pm 0.06$  fb<sup>-1</sup> [10].

We identify a reconstructed calorimeter cluster as a photon when it satisfies the following requirements (photon ID): (i) at least 90% of the energy is deposited in the EM section of the calorimeter; (ii) the calorimeter isolation variable  $\mathcal{I} = [E_{\text{tot}}(0.4) - E_{\text{em}}(0.2)]/E_{\text{em}}(0.2)$  is less than 0.07, where  $E_{\text{tot}}(0.4)$  denotes the total energy deposited in the calorimeter in a cone of radius  $\mathcal{R} = \sqrt{(\Delta\eta)^2 + (\Delta\phi)^2} = 0.4$ , and  $E_{\text{em}}(0.2)$  is the EM energy in a cone of radius  $\mathcal{R} = 0.2$ ; (iii) the track isolation variable, defined as the scalar sum of the transverse momenta ( $p_T$ ) of all tracks which originate from the interaction vertex in an annulus of  $0.05 < \mathcal{R} < 0.4$  around the cluster, is less than 2 GeV; (iv) it is in the CC with  $|\eta| < 1.1$ ; (v) both transverse and longitudinal shower shapes are consistent with those of a photon; (vi) it has neither an associated track in the central tracking system nor a significant density of hits in the SMT and CFT systems consistent with the presence of a track with  $p_T$  in agreement with its transverse energy; and (vii) there is energy deposit in the CPS matched to it. Jets are reconstructed using the iterative midpoint cone algorithm [11] with a cone size of 0.5. The missing transverse energy is computed from calorimeter cells with  $|\eta| < 4$  and corrected for the EM and jet energy scales.

The *photon* sample is obtained by selecting events with only one photon with  $p_T > 90$  GeV, at least one reconstructed interaction vertex consistent with the measured direction of the photon (see below), and  $\cancel{E}_T > 70$  GeV. This high  $\cancel{E}_T$  requirement guarantees negligible multijet background in the final candidate sample. Additionally, in order to avoid large  $\cancel{E}_T$  due to mismeasurement of jet energy, we require no jets with  $p_T > 15$  GeV. We reject events with reconstructed muons and with cosmic ray muons identified by the muon scintillator counters timing signal or by the presence of a characteristic pattern of hits in the muon drift chambers aligned with the reconstructed photon. In order to further reject events with leptons that leave a distinguishable signature in the tracker but that are not reconstructed in the other subsystems of the detector, we impose a requirement on the

transverse momentum of any isolated track not to be greater than 6.5 GeV. A track is considered to be isolated if the ratio between the scalar sum of the  $p_T$  of all tracks which originate from the interaction vertex in an annulus of  $0.1 < \mathcal{R} < 0.4$  around the track and the  $p_T$  of the track is less than 0.3.

The EM pointing algorithm allows the calculation of the direction of the EM shower based on the transverse and longitudinal segmentation of the calorimeter and preshower systems. EM pointing is performed independently in the azimuthal and polar planes. The former results in the measurement of the distance of closest approach to the beam line (DCA), and the latter in the prediction of the  $z$  position (along the beam) of the interaction vertex in the event. We require that the  $z$  coordinate of at least one interaction vertex in the event be within 10 cm of the position predicted by the pointing algorithm, and use the DCA to estimate the remaining backgrounds from jet-photon misidentification and non-collision muons. Misidentified jets have poor pointing resolution, and therefore (compared to electrons or photons) a wider DCA distribution is expected. Likewise, one can anticipate the DCA distribution for non-collision events to have an even wider shape. After these requirements, 35 events are selected in the *photon* sample.

We prepare three DCA distribution templates: the *non-collision* template, the *misidentified jets* template, and the  *$e/\gamma$*  template. The first template is obtained by selecting events with no hard scatter (no reconstructed interaction vertex or total number of reconstructed tracks less than three), or events with identified cosmic muons. The *misidentified jets* template is extracted from the *fake photon* sample, which fulfills exactly the same requirements as the *photon* sample except that the photon track isolation requirement is inverted. This sample is dominated by misidentified jets. Finally, the  *$e/\gamma$*  template is obtained from a real data sample of isolated electrons.

The total number of background events from misidentified jets ( $N_{\text{fake}}$ ) can be predicted from the *fake photon* sample if one knows the rates at which jets, passing all other photon ID criteria, fail or pass the track isolation requirement. To measure those rates we use an *EM plus jet* sample, where the EM object passes all photon ID requirements except the track isolation, and where the jet approximately balances the EM object in the transverse plane. We first determine the number of events ( $N_1$ ) in the sample that fail the track isolation requirement. We then examine the DCA distribution for events that pass the track isolation, and exploit the difference in DCA distributions of real and misidentified photons to extract the number of the latter passing the track isolation ( $N_2$ ).  $N_{\text{fake}}$  is then equal to the number of events in the *fake photon* sample multiplied by  $N_2/N_1$ . We fit the *photon* sample DCA distribution to a linear sum of the three templates, fixing the contribution of *misidentified jets* as described above, and determine the  *$e/\gamma$*  and *non-*

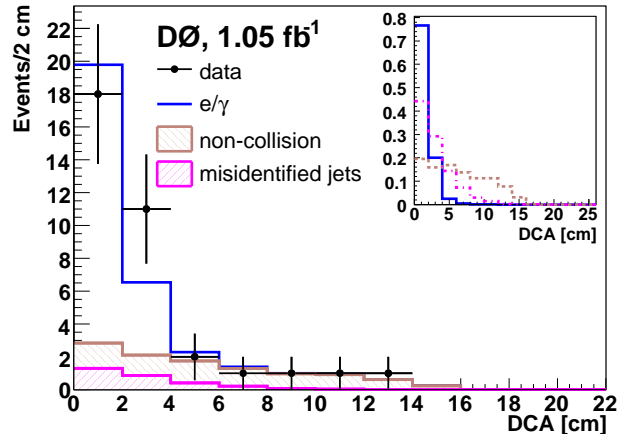


FIG. 1: Template fit in DCA of the selected events in data. Inset plot displays individual template shapes.

*collision* contributions. The result of the fit is illustrated in Fig. 1. Most of the signal photons are within 4 cm in DCA, therefore we limit our analysis to this particular window which contains 29 data events.

The only physics background to the  $\gamma + \cancel{E}_T$  final state comes from the process  $Z + \gamma \rightarrow \nu \bar{\nu} + \gamma$ . This irreducible contribution is estimated from a sample generated with PYTHIA [12] using CTEQ6L1 parton distribution functions (PDF) [13]. The main instrumental background arises from  $W \rightarrow e\nu$  decays, where the electron, due to tracking inefficiency or hard bremsstrahlung, is misidentified as a photon. This contribution is estimated from data using a sample of isolated electrons. The same requirements as for the *photon* sample are imposed, and the remaining number of events is scaled by  $(1 - \epsilon_{\text{trk}})/\epsilon_{\text{trk}}$ , where  $\epsilon_{\text{trk}}$  is the track reconstruction efficiency determined to be  $(98.6 \pm 0.1)\%$  [14]. A smaller instrumental contribution to the background is expected from  $W + \gamma$  production where the charged lepton in a leptonic  $W$  boson decay is not detected. The kinematics of this contribution is obtained from  $W(+\text{jets}) \rightarrow \text{lepton} + \nu(+\text{jets})$  Monte Carlo (MC) samples generated with PYTHIA, while the cross section is taken from the MC generator based on [15], which predicts all contributions (initial state radiation, trilinear gauge boson vertex, and final state radiation) to the full process. We generate signal events [16] with  $M_D = 1.5$  TeV for  $n = 2, 3, 4, 5, 6, 7$  and 8. For different values of  $M_D$ , the cross section scales as  $1/M_D^{n+2}$ , leaving the kinematic spectra unaffected for a fixed number of extra dimensions.

All MC events are passed through a detector simulation based on the GEANT [17] package, and processed using the same reconstruction software as for the data. Additionally, we apply scale factors which account for the differences between the efficiency determinations from data and simulation.

TABLE I: Data and estimated backgrounds.

Background	Number of expected events
$Z + \gamma \rightarrow \nu\bar{\nu} + \gamma$	$12.1 \pm 1.3$
$W \rightarrow e\nu$	$3.8 \pm 0.3$
Non-collision	$2.8 \pm 1.4$
Misidentified jets	$2.2 \pm 1.5$
$W + \gamma$	$1.5 \pm 0.2$
Total Background	$22.4 \pm 2.5$
Data	29

The main sources of systematic error are the uncertainty in the photon identification efficiency (5%), the uncertainty in the total integrated luminosity (6.1%), and 4% uncertainty from the PDF. For the case of SM backgrounds estimated from MC, the quoted errors include the uncertainty on the cross sections, which is dominated by the uncertainty in the k-factors (7%) [15, 18]. The uncertainty in the width of the  $e/\gamma$  sample DCA template results in an additional systematic error of 0.4 events in the non-collision background estimate.

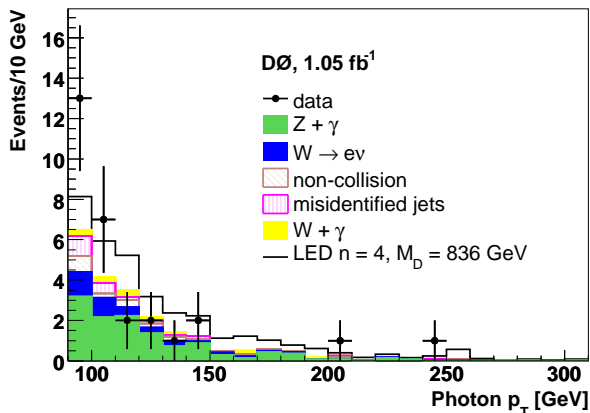


FIG. 2: Photon  $p_T$  distribution for the final candidate events, after all the applied requirements. The LED signal is stacked on top of SM backgrounds.

The final number of events for data and backgrounds are given in Table I. Fig. 2 shows the photon  $p_T$  distribution, with the SM backgrounds stacked on top of each other. Data and the SM expectation agree, so we proceed to set lower limits for the fundamental Planck scale  $M_D$ . We employ the Modified Frequentist approach [19] to set limits on the production cross section. This method is based on a log-likelihood ratio test statistic and uses the binned photon  $p_T$  distribution. At the 95% C.L. we find the following lower limits:  $M_D > 884, 864, 836, 820, 797, 797$  and  $778$  GeV for  $n = 2, 3, 4, 5, 6, 7$  and  $8$  extra dimensions, respectively. Table II and Fig. 3 summarize the results for the limit calculations.

To conclude, we have conducted a search for LED in

the  $\gamma + \cancel{E}_T$  channel, finding no evidence for their presence. We have set limits on the fundamental Planck scale, significantly improving results of previous searches.

TABLE II: Summary of limit calculations.

$n$	Signal efficiency	Observed (expected) cross section limit (fb)	Observed (expected) $M_D$ lower limit (GeV)
2	$0.49 \pm 0.04$	27.6 (23.4)	884 (921)
3	$0.48 \pm 0.04$	24.5 (22.7)	864 (877)
4	$0.47 \pm 0.04$	25.0 (22.8)	836 (848)
5	$0.43 \pm 0.04$	25.0 (24.8)	820 (821)
6	$0.50 \pm 0.05$	25.4 (22.3)	797 (810)
7	$0.49 \pm 0.04$	24.0 (23.1)	797 (801)
8	$0.52 \pm 0.05$	24.2 (21.9)	778 (786)

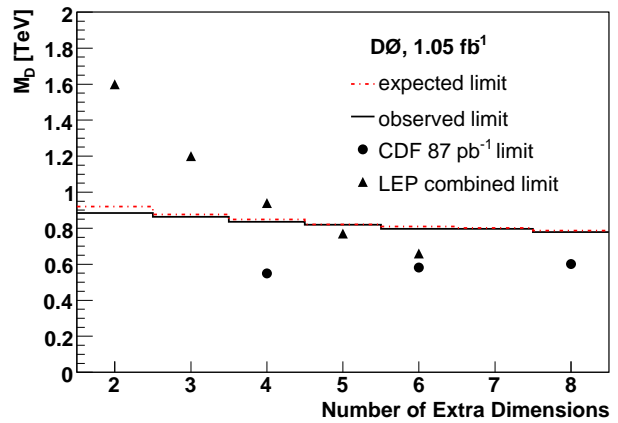


FIG. 3: Expected and observed lower limits on  $M_D$  for LED in the  $\gamma + \cancel{E}_T$  final state. CDF limits with  $87 \text{ pb}^{-1}$  of data [3], and LEP combined limits [6] are also shown. The data are the black points with statistical uncertainties.

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